Black Holes I — Exercise sheet 5

(5.1) Eddington-Finkelstein gauge

Take the Schwarzschild line-element in Schwarzschild gauge, viz.

$$ds^{2} = -\left(1 - \frac{2M}{r}\right) dt^{2} + \frac{dr^{2}}{\left(1 - \frac{2M}{r}\right)} + r^{2}\left(d\theta^{2} + \sin^{2}\theta d\phi^{2}\right)$$

and perform a coordinate transformation v = t + f(r) with some suitable function f (which you have to determine) so that you obtain the Schwarzschild line-element in Eddington-Finkelstein gauge, viz.

$$ds^{2} = -2 dr dv - \left(1 - \frac{2M}{r}\right) dv^{2} + r^{2} \left(d\theta^{2} + \sin^{2}\theta d\phi^{2}\right).$$

Is this line-element singular at r = 2M?

(5.2) Coordinate transformation of Christoffels

Consider a general coordinate transformation x' = x'(x) and calculate the transformation of the Christoffel symbols (of the first kind). Do they transform as a tensor?

(5.3) Covariant derivative of a constant vector field

Given a smooth manifold in $n \geq 2$ dimensions equipped with an arbitrary metric of signature $(-,+,\ldots,+)$ consider a vector field v^{μ} , whose components in a certain coordinate system (t,x_1,\ldots,x_{n-1}) are given by $v^t=1$ and $v^{\mu}=0$ otherwise. Thus, in these coordinates the vector field is a constant vector field and takes the simple form $v^{\mu}=(1,0,\ldots,0)$. The main task of this exercise is to calculate the covariant derivative of this vector field, $\nabla_{\nu}v^{\mu}$. (When) is it true that $\nabla_{\nu}v^{\mu}=0$ for this vector field v^{μ} ?

These exercises are due on November 10th 2015.

Hints:

- This is perhaps the most efficient way to proceed: use the Ansatz v = t + f(r), derive the coordinate differential dv and calculate backwards by starting with the line-element in Eddington-Finkelstein gauge and ending with the line-element in Schwarzschild gauge. Concerning the last question: the metric obviously is finite at r = 2M; the only issue not completely obvious is whether it is invertible at r = 2M.
- The definition of the Christoffel symbols of the first kind is

$$\Gamma_{\alpha\beta\gamma} = \frac{1}{2} \left(\partial_{\beta} g_{\alpha\gamma} + \partial_{\gamma} g_{\alpha\beta} - \partial_{\alpha} g_{\beta\gamma} \right)$$

Remember that the metric is a tensor and therefore transforms under coordinate change as any other tensor does.

This is a very short exercise (it should take less than a minute to write down the answers). However, be careful before answering the question

— it may be tempting to give a short but incorrect answer.